

Preresonance Raman Spectrum of the C₁₃H₉ Fluorene-like Radical

Jan Szczepanski, John Banisaukas, and Martin Vala*

Department of Chemistry and Center for Chemical Physics, University of Florida,
Gainesville, Florida 32611-7200

So Hirata and William R. Wiley

Environmental Molecular Sciences Laboratory, Pacific Northwest National Laboratory, P.O. Box 999,
Richland, Washington 99352

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The neutral open-shell species C₁₃H₉ formed from fluorene, C₁₃H₁₀, by low-energy electron bombardment and by ultraviolet photolysis in an argon matrix at 12 K has been studied via preresonance Raman, infrared, and ultraviolet/visible spectroscopy. Density functional theory calculations (B3LYP/6-31G(d,p)) of the CH bond energies of neutral fluorene showed that the most probable position for the hydrogen loss is the sp³ carbon in the five-membered ring. Calculations of the C₁₃H₉ harmonic vibrational frequencies are shown to match the experimental Raman (and infrared) bands well. A new electronic transition is identified at 283.1 nm (4.38 eV). Its position agrees with earlier time-dependent density functional theory calculations. Oscillator strengths for this transition and three others are estimated. The electronic transitions in the dehydrogenated species, C₁₃H₉, are strongly red-shifted compared to fluorene.

I. Introduction

The species responsible for the unidentified infrared (UIR) emission bands observed from many carbon-rich interstellar sources and the carriers of the diffuse interstellar visible absorption bands (DIBs) have occupied researchers for many years now.^{1–6} Polycyclic aromatic hydrocarbons (PAHs), in their neutral, ionized, hydrogenated, or dehydrogenated forms, have been proposed as possible carriers of both sets of bands. Since the PAHs were first proposed^{1,2} as the primary UIR emitters in 1984–5, the model has been much discussed and many laboratory experiments and astrophysical observations and calculations have been performed, which point to the viability of the idea. Until recently, there were no direct measurements that showed that PAH ions could emit in the infrared or that PAHs were present in the interstellar medium. But now Cernicharo et al. have reported⁷ infrared absorption bands in the vicinity of the protoplanetary nebula CRL 618, which they ascribe to the aromatic molecule, benzene. And, Saykally and co-workers have recently seen⁸ the infrared emission from a PAH cation, pyrene, in a laboratory experiment.

Dehydrogenated and fragmented PAHs have been receiving an increasing amount of attention. Using Fourier transform ion cyclotron resonance (FT-ICR) mass spectrometry, Ekern and co-workers reported the loss of up to five hydrogens from the fluorene cation (C₁₃H₁₀⁺)⁹ and the complete dehydrogenation of the coronene cation (C₂₄H₁₂⁺) and the naphthopyrene cation (C₂₄H₁₂⁺)¹⁰ after 1.5 s of UV (Xe lamp) photolysis. Recently, Dibben et al. demonstrated a similar photoprocess of stripping of H, C₂H₂, and C₄H₂ from fluorene cation to form pure C₉⁺ and C₁₁⁺ carbon clusters¹¹ in a proposed ring-opening mechanism.¹²

In this paper, we suggest that dehydrogenated PAHs may play a role in interstellar chemistry and could be responsible for some of the DIBs. It is general knowledge that, when ionized, neutral PAHs shift their electronic absorption bands to the red. But it is not widely known that removal of a hydrogen from a PAH may also result in a red shift. In general, lower-energy photons are required for hydrogen removal than for ionization of neutral PAHs. Thus, dehydrogenation could easily compete with ionization processes in regions of the interstellar medium (ISM) where penetration of low-temperature radiation is important. From recent laboratory experiments, it is now known that the photodehydrogenation of some PAH cations can occur very efficiently. Dehydrogenated neutral and ionic PAHs may therefore play a crucial role in the photochemistry of the ISM.

In this work, the spectral characterization of a neutral dehydrogenated PAH, C₁₃H₉, is presented. The C₁₃H₉ species was produced from fluorene (C₁₃H₁₀) by removal of one hydrogen (most likely from the sp³ carbon) by either UV photolysis or low-energy electron bombardment. The Raman active modes were resonantly excited by 496.5 nm laser radiation, which corresponds to the low-energy absorption band shoulder of the ²A₂(D₂) ← ²B₂(D₀) (0-0) transition of C₁₃H₉. This electronic transition was recently identified from a combination of experimental observations and vertical excitation energy calculations using time-dependent density functional theory (TDDFT).¹³ The results of density functional theory (B3LYP/6-31G(d,p)) calculations of the harmonic frequencies of Raman and IR modes are also reported and compared to experimental data.

II. Computational Methods

Geometry optimization and calculation of harmonic frequencies for the C₁₃H₉ radical were carried out using density functional theory (DFT) employing the Becke3–Lee–Yang–Parr (B3LYP) functional with the 6-31G(d,p) basis set using

* To whom correspondence should be addressed. E-mail: mvala@chem.ufl.edu.

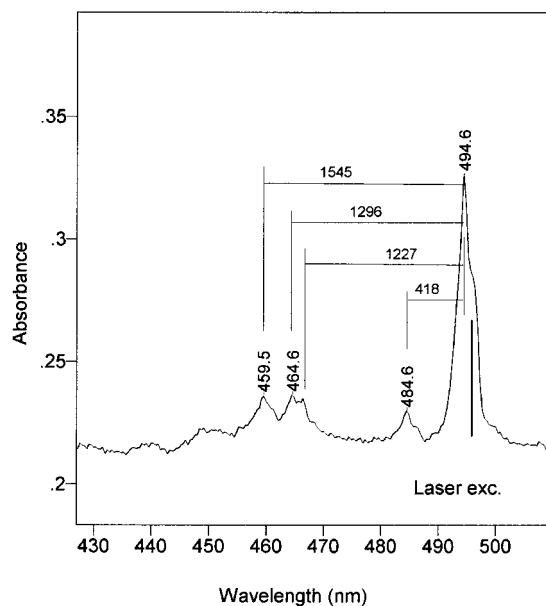


Figure 2. The ${}^2A_2(D_2) \leftarrow {}^2B_2(D_0)$ electronic absorption spectrum of the C₁₃H₉ radical isolated in Ar matrix at 12 K. The spectrum was recorded after 4 h matrix deposition during low-energy electron bombardment of the C₁₃H₁₀/Ar beam. The 496.5 nm Ar⁺ laser line was used to resonantly excite the Raman vibrations (cf. Figure 3). Some vibrational mode frequencies (in cm⁻¹) active in the D₂ state are marked in the spectrum as well. Note that for the neutral C₁₃H₉ radical (an open-shell system) the lowest-energy spin- and symmetry-allowed electronic transition is red-shifted into the visible energy region from the ultraviolet region of the first electronic absorption of the parent fluorene.

C–H bonds in this species. A similar effect was observed earlier for the ionic system C₁₃H₁₀⁺/C₁₃H₉⁺, in which the C–H₉ bond in C₁₃H₉⁺ was not broken in a number of different photofragmentation processes.^{12,13} The potential energy curve along the C–H₉ stretching coordinate was computed in the C₁₃H₁₀ ground electronic state to determine whether any energy barrier to the abstraction of the H₉ atom was present. None was found. No barrier was found for the similar abstraction in the C₁₃H₁₀⁺ cation either.

B. Preresonance Raman Spectrum. Because preresonance and resonance Raman spectra are affected only by the electronic ground-state vibrational modes, the ${}^2B_2(D_0)$ state harmonic mode frequencies were calculated (B3LYP/6-31G(d,p)) for the C₁₃H₉ species. These values, as well as the calculated Raman intensities, are compared to observed preresonance Raman bands in Table 1. The band intensity distribution is expected to be different in the preresonance Raman spectrum compared to the normal Raman distribution because of the intensity enhancement effect for the modes involved in the electronic transition. In fact, the 428 and 1575 cm⁻¹ D₀ state modes (cf. Table 1 and Figure 3) are also observed in the electronic absorption to D₂ as 418 and 1545 cm⁻¹ (cf. Figure 2), respectively. In the preresonance Raman spectrum, these modes increase their relative intensities from 0.02 and 0.01 (as predicted for normal Raman, cf. Table 1) to 1.0 and 0.9, respectively.

As usual with resonantly enhanced Raman modes, the overtone and combination modes observed in the spectrum of Figure 3 also have relatively large intensities. These involve the low-energy a₁ symmetry modes, 222, 428, and 883 cm⁻¹ (cf. Table 1). The expected symmetry of the overtone and combination modes are a₁, because the direct product of a₁ × a₁ = a₁ for C_{2v} symmetry. The intensity enhancement is

TABLE 1: Preresonance Raman Band Assignments of Fluorene-like C₁₃H₉ Radical Observed in Solid Ar at 12 K (cf. Figure 3)^a

mode assignment ^b	mode frequencies		intensities	
	obsd, cm ⁻¹	calcd, cm ⁻¹	obsd, rel	calcd, Å ⁴ amu ⁻¹
$\tau(\text{CCC}) + \epsilon(\text{CH}), a_1$	222	212	0.13	4 (0.02)
$\tau(\text{CCC}) + \beta(\text{CH}), a_1$	429 [418]	419	1.00	13 (0.05)
222 + 429	650		0.90	
$\epsilon(\text{CH}) + \tau(\text{CCC}), b_1$	678	696	0.20	2 (0.01)
$\epsilon(\text{CH}) + \tau(\text{CCC}), a_2$		742		4 (0.02)
2 × 429	849		0.16	
$\epsilon(\text{CH}) + \tau(\text{CCC})b_1$		859		8 (0.04)
$\epsilon(\text{CH}) + \tau(\text{CCC})a_1$	883	864	0.27	31 (0.12)
$\epsilon(\text{CH}) + \tau(\text{CCC})a_2$		919		2 (0.02)
$\epsilon(\text{CH}) + \tau(\text{CCC})b_2$		986		11 (0.04)
$\beta(\text{CH}) + \text{R}(\text{CC}), a_1$	1021	1020	0.07	65 (0.25)
222 + 2 × 429	1071		0.16	
$\beta(\text{CH}) + \text{R}(\text{CC}), b_2$		1092		13 (0.05)
222 + 883	1104		0.20	
$\beta(\text{CH}) + \text{R}(\text{CC}), a_1$	1120	1096	0.95	23 (0.09)
$\beta(\text{CH}) + \text{R}(\text{CC}), a_1$	1163	1156	0.40	25 (0.10)
$\beta(\text{CH}) + \text{R}(\text{CC}), a_1$	1194 [1227]	1192	0.90	227 (0.87)
$\beta(\text{CH}) + \text{R}(\text{CC}), b_2$	1215	1218	0.31	28 (0.10)
$\beta(\text{CH}) + \text{R}(\text{CC}), a_1$	1256 [1296]	1249	0.87	262 (1.00)
$\beta(\text{CH}) + \text{R}(\text{CC}), a_1$		1296		37 (0.14)
$\text{R}(\text{CC}) + \beta(\text{CH}), b_2$		1340		39 (0.15)
$\text{R}(\text{CC}), a_1$	1368	1370	0.41	41 (0.16)
$\text{R}(\text{CC}), a_1$	1443	1442	0.29	90 (0.34)
$\text{R}(\text{CC}), b_2$		1479		3 (0.01)
$\text{R}(\text{CC}), a_1$		1480		44 (0.17)
$\text{R}(\text{CC}), b_2$		1580		15 (0.06)
$\text{R}(\text{CC}), a_1$	1575 [1545]	1582	0.90	3 (0.01)
$\text{R}(\text{CC}), b_2$		1589		100 (0.38)

^a The calculated Raman harmonic frequencies at B3LYP/6-31G(d,p) level for ${}^2B_2(D_0)$ ground state of C₁₃H₉ are scaled uniformly by 0.978 factor. All fundamental mode frequencies (in cm⁻¹) in the 100–1650 cm⁻¹ range with Raman intensities equal to 2 Å⁴/amu and larger are listed. The observed mode frequencies (in cm⁻¹) in the ${}^2A_2(D_2)$ electronic excited state are given in brackets, while the relative intensities are in parentheses. Note that the calculated normal Raman mode intensities are only loosely connected to the preresonance Raman band intensities (see text). ^b Notation used: R are stretching vibration modes, α and β are in-plane bending vibration modes, and ϵ and τ are out-of-plane vibration modes. The first vibration indicated, in modes of mixed character, contributes the largest amount to the total vibration. The mode symmetries in C_{2v} point group are given in the second column.

observed for the 678, 1120, 1163, 1215, and 1368 cm⁻¹ fundamental modes also.

Geometry optimization (B3LYP/6-31G(d,p)) reveals that the fluorene skeleton is preserved in the C₁₃H₉ species. Because both C₁₃H₉ and C₁₃H₁₀ have the same symmetry (C_{2v}), similar mode frequencies and symmetries are expected. This expectation is realized. The Raman depolarization ratios, ρ , of the 221, 421, 846, 1022, 1093, 1156, 1195, 1238, 1349, 1480, and 1578 cm⁻¹ bands determined by Bree and Zwarich for single-crystal fluorene lie in the range 0 < ρ < 0.75, indicating that they all have a₁ symmetry.¹⁹ These frequencies are in concert with the 222, 429, 883, 1021, 1120, 1163, 1194, 1256, 1368, 1443, and 1575 cm⁻¹ frequencies observed here (cf. Figure 3). All are here assigned to a₁ symmetry (cf. Table 1). Two bands were observed in the Raman spectrum (cf. Figure 3) with mode symmetries different than a₁; one is of b₁ symmetry at 678 cm⁻¹, in close coincidence to the band calculated at 696 cm⁻¹, also of b₁ symmetry. The second one is of b₂ symmetry at 1215 cm⁻¹ for which the predicted (B3LYP/6-31G(d,p)) band lies at 1218 cm⁻¹ (b₂ symmetry, cf. Table 1). In the single-crystal

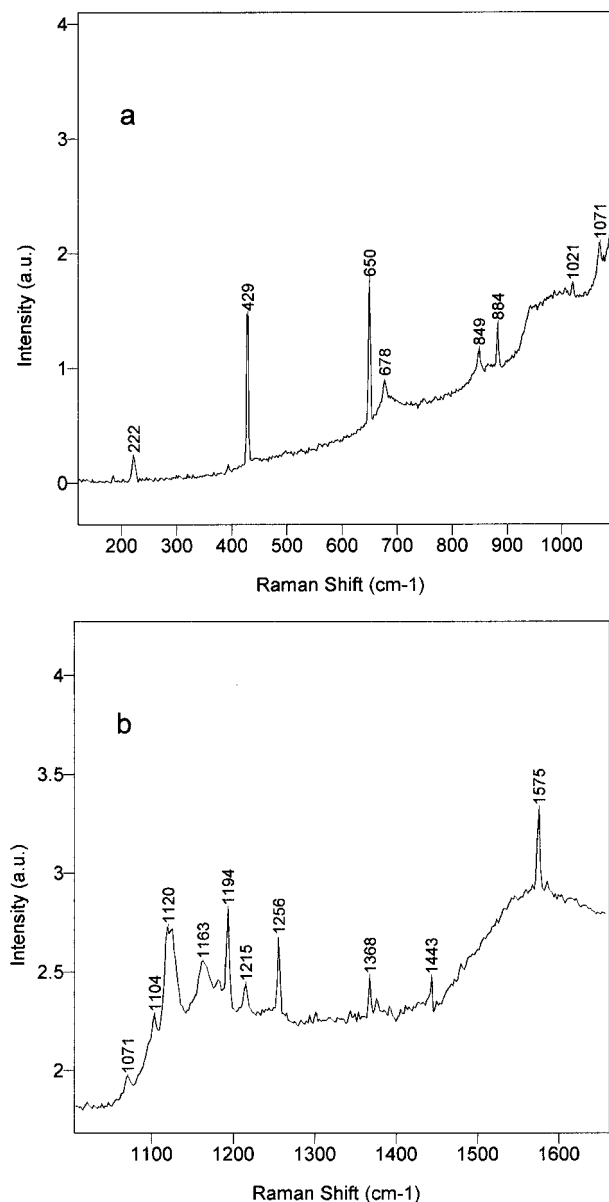


Figure 3. The preresonance Raman spectrum of $C_{13}H_9$ fluorene-like radical trapped in solid Ar at 12 K in the (a) 100–1100 and (b) 1000–1650 cm^{-1} regions. Laser excitation at 496.5 nm (50 mW) was used (cf. Figure 2). The Raman band shifts (in cm^{-1}) are marked. No resonantly enhanced Raman modes over 1600 cm^{-1} were observed (see text for the details).

Raman spectrum, two bands, 1193 and 1172 cm^{-1} , both of b_2 symmetry, were identified in this region.¹⁹

The preresonance Raman spectrum of the $C_{13}H_9$ species was scanned to 3200 cm^{-1} , but no bands in the region greater than 1600 cm^{-1} were observed. This is typical of the resonance Raman of aromatic species in which mostly resonantly enhanced ring modes but no C–H stretches are observed, even though in normal Raman spectra, these C–H modes are very strong. The resonance (or preresonance) Raman spectra of *p*-dichlorobenzene radical cation²⁰ and naphthalene radical cation²¹ are good examples of this.

C. Infrared Absorption Spectrum. The IR and UV–visible absorption spectra were collected on the same $C_{13}H_9/C_{13}H_{10}/Ar$ matrix. The infrared spectrum of the $C_{13}H_9$ radical revealed, in addition to the 722.2 and 780.9 cm^{-1} bands reported earlier,¹³ new bands at 1239.5 and 1567.8 cm^{-1} . The intensities of the 1239.5 and 1567.8 cm^{-1} bands tracked the intensities of the

former two bands under all experimental conditions used (i.e., EI with different electron beam energies and UVP with different photon fluxes). Intensities ratios of 1.0:0.98:0.10:0.43 for the 722.2, 780.9, 1239.5, and 1567.8 cm^{-1} bands were observed, respectively. From calculations, B3LYP/6-31G(d,p) (B3LYP/6-31++G(d,p)), harmonic IR frequencies (scaled by 0.978) were found at 725 [52 km/mol] (723 [94 km/mol]), 778 [44 km/mol] (785 [48 km/mol]), 1248 [10 km/mol] (1245 [10 km/mol]), and 1580 cm^{-1} [27 km/mol] (1572 cm^{-1} [31 km/mol]), with intensity ratios of 1.0(1.0):0.87(0.51):0.19(0.11):0.51(0.33), respectively. The agreement with experiment is very good.

D. Electronic Absorption Spectrum. A portion of the electronic absorption spectrum of the fluorene-like $C_{13}H_9$ radical was reported recently.¹³ TDDFT (B3LYP/6-31(2+,2+)G(d,p)) calculations¹³ predicted a ${}^2A_2 \leftarrow {}^2B_2(D_0)$ electronic transition in $C_{13}H_9$ at 4.37 eV ($f_{TDDFT} = 0.0154$) that was not observed in our earlier work. An extended search for such a band in a series of experiments in the UV–vis–IR regions (on the same $C_{13}H_9/Ar$ matrices) revealed a candidate at 283.1 nm (4.38 eV). This band (cf. Figure 4b) displayed parallel intensity behavior with the previously assigned 494.6 (2.51), 370.6 (3.34), and 353.2 nm (3.51 eV) bands, as well as the 722.2, 780.9, 1239.5, and 1580 cm^{-1} infrared bands, all attributed to $C_{13}H_9$ radical. It is concluded that the 283.1 nm band is the predicted ${}^2A_2 \leftarrow {}^2B_2(D_0)$ 0-0 transition in $C_{13}H_9$.

Because the electronic and IR absorption spectra were recorded on the same sample/matrices in the above experiments, it is possible to determine the oscillator strengths of the optical bands. The experimental ratio of the ${}^2A_2(D_2) \leftarrow {}^2B_2(D_0)$ (0-0) absorption transition (494.6 nm) and the 780.9 cm^{-1} absorption (well-isolated IR band, not displayed) was determined as $A(494.6)/A(780.9) = 43.6$. Assuming that the B3LYP/6-31G(d,p) integral intensity of 44 km/mol = $2.303\epsilon_{max}(780.9)\Delta\nu_{1/2}$ for the 778 cm^{-1} (780.9 cm^{-1} experimental) band is reliable, the molar absorption coefficient, $\epsilon_{max}(494.6)$, of the 494.6 nm $C_{13}H_9$ band is found equal to 33 340 $M^{-1} cm^{-1}$. From this value, an oscillator strength, $f_{est} = 0.025$, was estimated for the experimental absorption spectrum in Figure 2. With the use of the B3LYP functional with the diffuse functions added to the basis set (B3LYP/6-31++G(d,p)), a 10% increase in the integral intensity of the calculated 785 cm^{-1} band was observed. This increase reflects a similar increase for all f_{est} values in Table 2, when B3LYP/6-31++G(d,p) is applied in the oscillator strength estimated. The spectrum was assigned¹³ to the ${}^2A_2(D_2) \leftarrow {}^2B_2(D_0)$ transition of $C_{13}H_9$ from TDDFT calculations, which predicted an oscillator strength of $f_{TDDFT} = 0.0525$. Using similar procedures, we also estimated the oscillator strengths of the other absorption bands in the Figure 4. The values are given in Table 2.

V. Conclusions

(1) On the basis of the calculated low energy (3.38 eV) for the C–H₉ bond in $C_{13}H_{10}$, the first dehydrogenation in fluorene is most likely the loss of H₉ from the sp^3 carbon.

(2) The preresonance Raman spectrum of the fluorene-like $C_{13}H_9$ radical was recorded and analyzed. Density functional theory calculations (B3LYP/6-31G(d,p)) of the harmonic mode frequencies are shown to agree well (after scaling) with the observed band positions. Vibrational mode symmetries agree with the assignments of fundamentals reported for single-crystal fluorene.

(3) Two new infrared bands at 1239.5 and 1567.8 cm^{-1} , together with the previously observed 722.2 and 780.9 cm^{-1} bands, were identified as arising from the $C_{13}H_9$ radical in solid Ar.

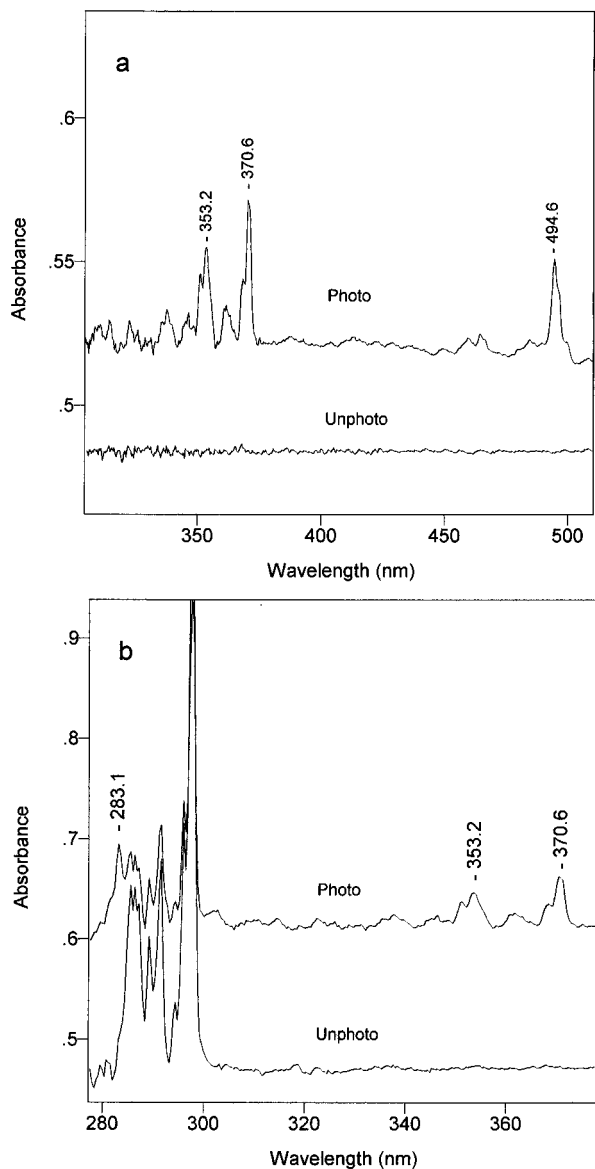


Figure 4. Part of the electronic absorption spectrum of fluorene neutral, C₁₃H₁₀, isolated in solid Ar at 12 K (before UV photolysis, lower spectrum) and the absorption spectrum of the mixture of C₁₃H₁₀ and C₁₃H₉ fluorene-like radical in solid Ar (recorded after 2 h of photolysis of C₁₃H₁₀/Ar matrix) displayed in two energy regions: (a) 340–500 nm; (b) 280–370 nm. The marked bands, attributed to the C₁₃H₉ radical, are assigned in Table 2. The strong band at 297.5 nm is due to the S₁ ← S₀ (0-0) transition of parent fluorene. These spectra are baseline-corrected to cancel the steadily rising signal with decreasing wavelength due to Rayleigh scattering. Note the ca. 50% intensity decline of the parent fluorene bands after UV photolysis.

(4) The newly observed UV band at 283.1 nm was tentatively assigned to a ${}^2A_2 \leftarrow {}^2B_2(D_0)$ transition of the C₁₃H₉ radical on the basis of the energy coincidence with the calculated vertical excitation energies¹³ as well as of band intensity correlations with the known bands of C₁₃H₉. From the intensity-correlated IR and UV–visible bands (recorded on the same C₁₃H₉/Ar matrix) and from the calculated (B3LYP/6-31G(d,p)) IR integral band intensities, the oscillator strengths, f_{est} , for the observed UV–visible bands were estimated and shown to compare reasonably well with the calculated f_{TDDFT} values.

(5) The preresonance Raman spectrum obtained with 496.5 nm laser excitation combined with calculated Raman frequencies of C₁₃H₉ confirms the earlier assignment of the ${}^2A_2(D_2) \leftarrow$

TABLE 2: Oscillator Strength Values, f_{est} , Estimated from the Intensity-Correlated IR and UV–Visible Bands (recorded on the same C₁₃H₉/Ar Matrix) and from the Calculated (B3LYP/6-31G(d, p)) IR Band Integral Intensities^a

state ^b	calcd ^b		expt	
	energy, eV	f_{TDDFT}	energy, eV (nm)	f_{est} ^c
${}^2A_2 (\pi_0 \leftarrow \pi_{-2})$	2.75	0.0525	2.51 (494.6) ^b	0.025
${}^2B_2 (\pi_0 \leftarrow \pi_{-4})$	3.69	0.0153	3.34 (370.6) ^b	>0.050 ^d
${}^2A_2 (\pi_2 \leftarrow \pi_0)$	3.93	0.0359	3.51 (353.2) ^b	< 0.037 ^d
${}^2A_2 (\pi_1 \leftarrow \pi_{-2})$	4.37	0.0154	4.38 (283.1) ^c	>0.012 ^e

^a The calculated f_{TDDFT} oscillator strengths for fluorene-like C₁₃H₉ radical using TDDFT with B3LYP functional and 6-31(2+,2+)(G(d,p) basis set are given for the sake of comparison. TDDFT vertical excitation energies (eV) and experimental 0-0 absorption band positions (eV (nm)) from the 2B_2 ground state are listed. ^b From ref 13. ^c This work. The f_{est} values are 10% higher when the B3LYP/6-311++G(d,p) level is applied. ^d The uncertain f_{est} value is due to partial overlap with the ${}^2B_2 \leftarrow {}^2B_2(D_0)$ or ${}^2A_2 \leftarrow {}^2B_2(D_0)$ transition (cf. Figure 4a). ^e f_{est} value error is due to the lack of observation of high-energy vibrational progression bands for this transition and blending with the parent fluorene S₁ ← S₀ absorption bands (cf. Figure 4b).

${}^2B_2(D_0)$ (0-0) absorption transition at 494.6 nm (in solid Ar) to the C₁₃H₉ fluorene-like radical.

It is well-known that upon ionization the electronic absorption spectra of many PAHs undergo a considerable red-shift. We have shown here that a similar red-shift occurs in the neutral open-shell system C₁₃H₉ upon single hydrogen loss from the parent fluorene. The lowest-energy spin- and symmetry-allowed electronic transition is shifted from the ultraviolet into the visible region after the loss of one hydrogen. This is apparently typical for small and medium-sized closed-shell PAH neutrals. Radziszewski observed that dehydrogenated neutral benzene radical shifted its lowest electronic absorption into the visible region.²² Because PAH cations possess strong electronic absorption bands in the visible region, they were originally proposed as possible carriers of some of the diffuse interstellar bands (DIBs).^{23–25} The DIBs are a series of over 200 absorption bands of varying bandwidth observed from many regions of interstellar space where high carbon abundances are expected. If the observation reported here, that dehydrogenation of neutral closed-shell PAHs yielding neutral open-shell systems exhibiting bands red-shifted into the visible, can be shown to be a general phenomenon, these systems may then need to also be considered carefully as possible carriers of the DIBs.

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